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Physics-Informed Neural Networks Applied to a Series of Constrained Space Guidance Problems

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Introduction

- Optimal Control Problems (OCPs) represent an essential field within space engineering.
 - It is important to design minimum time, fuel/energy optimal trajectories for space missions.
- The goal of this work is to present a new methodology to solve Constrained Optimal Control Problems (COCPs) for space guidance by means of the novel Physics-Informed Neural Network (PINN) framework named Extreme Theory of Functional Connections (X-TFC).
 - Indirect method exploiting the Pontryagin Minimum Principle (PMP) is used to retrieve the optimal control.
 - Problems considered: Feldbaum problem (typical OCP), minimum time energy optimal Halo Halo transfer, 1D fuel optimal lunar landing.

Constrained Optimal Control Problems (COCPs)

Constrained Optimal Control Problem (COCP)

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Physics-Informed Neural Networks

- (Data) + Neural Networks + Physics Laws = Physics-Informed Neural Networks (PINN)
- PINNs are a newly developed framework for solving parametric DEs
 - The physics laws (modeled via parametric DEs), and eventually data, drive the training of the network



Image taken from: Lu, L., Meng, X., Mao, Z. and Karniadakis, G.E., 2019. DeepXDE: A deep learning library for solving differential equations. *arXiv preprint arXiv:1907.04502*.

PINN and TFC

- The **Theory of Functional Connections (TFC)** [Mortari, 2017] is a recently developed framework for functional interpolation
 - The functions are approximated via a **constrained expression**
 - Sum of a **free-chosen** function and a functional that **analytically** satisfies the constraints
 - TFC can be applied to solve DEs
 - The free-chosen function is an expansion of Chebyshev polynomials
 - The constraints are the Initial/Boundary Conditions (IC or BC)
- The **Physics-Informed Neural Network (PINN) Methods** are a novel approach, coming from the Machine Learning community
 - The DE latent solutions are approximated via a (Deep) Neural Network (NN), and the DEs drive the NN training (i.e., it acts as regulator)



Extreme Theory of Functional Connections (X-TFC)

• The Physics-Informed **Extreme Theory of Functional Connections (X-TFC)** is a synergy of the TFC and the standard PINN methods that helps to overcome their limitations for solving DEs.



Extreme Theory of Functional Connections (X-TFC)

• X-TFC uses the TFC constrained expression where the free-chosen function *g* is a Single Layer Feedforward NN (SLNN) trained via Extreme Learning Machine (ELM) Algorithm [Huang et al., 2006].



Extreme Theory of Functional Connections (X-TFC)



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• The Jacobian matrix required for an eventual iterative least-square procedure can be computed either analytically, or by the symbolic computation, or the automatic differentiation toolbox. 8/21

Extreme Learning Machine (ELM)

- ELM is a training algorithm for Single Layer Feedforward Neural Network (SLNN) that randomly selects input weights and bias, and computes the output weights (β) via least-square.
 - Input weights (w_i) and bias (b_i) are not tuned during the training.
- The convergence of the ELM algorithm is proved by Huang et al (2006).
 - The convergence is guaranteed for any input weights and bias randomly chosen from any continuous probability distribution.



SLNN example

• The Feldbaum problem is a typical optimal control problem.

Modified UOCP



• The following transversality condition has to be applied: $\lambda(1) = \lambda_f = 0$.

• The CEs and their derivatives are:

- Ω are the switching functions and their analytical expressions are computed by imposing the boundary conditions.
- The unknowns of the problem are: $\boldsymbol{\beta} = \{ \boldsymbol{\beta}_f \ \boldsymbol{\beta}_{\lambda} \ \boldsymbol{\beta}_w \ \boldsymbol{\beta}_u \ \boldsymbol{\beta}_{\nu} \}^{\mathrm{T}}$
- The losses of the associated Two-Point Boundary Value Problem (TPBVP) are:

$$\mathcal{L}_{f} = \dot{f} + f - u$$

$$\mathcal{L}_{\lambda} = \dot{\lambda} - \lambda + f$$

$$\mathcal{L}_{u} = u + \lambda + \nu$$

$$\mathcal{L}_{w} = 2\epsilon w - \nu \phi'(w)$$

$$\mathcal{L}_{d} = u - \phi(w)$$
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• Parameters employed for the Feldbaum problem:

n	L	σ	$[u_{min}, u_{max}]$	c
100	25	Gaussian	[-0.2,0]	4

- For all the chosen problems, the **initial guess** of the unknowns are chosen **randomly** within the interval (0,1).
- In order to make the UOCP close enough to the original COCP, a continuation procedure has been applied to decrease the value of ϵ from 1 to 10⁻⁵, while maintaining a good accuracy.
- Decreasing the value of ϵ induces higher values of the fictitious control w, which approaches the asymptotic limits.



• Results for the constrained Feldbaum problem with $\epsilon = 10^{-5}$



Trajectory and control plus the comparison with the unconstrained Feldbaum problem version

Performances of the constrained Feldbaum problem

- The Circular Restricted Three Body Problem (CR3BP) framework is employed to study **constrained minimum time energy optimal Halo-Halo transfers** in the Earth-Moon system (dimensionless units are considered).
- Due to the complexity of the problem, a **2-loop optimization procedure** is employed. An outer loop based on the Particle Swarm Optimization (PSO) is used to minimize the time, whereas the fixed time energy optimal problem is solved rapidly via indirect method and X-TFC in the inner loop.

$$\mathcal{J} = \Gamma t_f + \frac{1}{2} \int_{t_0}^{t_f} \mathbf{u}^T \mathbf{u} \, dt + \epsilon \int_{t_0}^{t_f} \mathbf{w}^T \mathbf{w} \, dt \qquad \text{subject to:} \qquad \begin{cases} \dot{\mathbf{r}} = \mathbf{v} \\ \dot{\mathbf{v}} = \nabla U(\mathbf{r}) + \mathbf{M} \mathbf{v} + t_0 \leq t \leq t_f \\ \mathbf{r}(t_0) = \mathbf{r}_0 \\ \mathbf{v}(t_0) = \mathbf{v}_0 \\ \mathbf{v}(t_0) = \mathbf{v}_0 \\ \mathbf{r}(t_f) = \mathbf{r}_f \\ \mathbf{v}(t_f) = \mathbf{v}_f \\ u_{\min} \leq u_i \leq u_{\max} \end{cases}$$

• The Hamiltonian of the problem is:

min

$$H = \frac{1}{2}\mathbf{u}^T\mathbf{u} + \epsilon\mathbf{w}^T\mathbf{w} + \boldsymbol{\lambda}_r^T\mathbf{v} + \boldsymbol{\lambda}_v^T(\nabla U(\mathbf{r}) + \mathbf{M}\mathbf{v} + \mathbf{u}) + \boldsymbol{\nu}^T(\mathbf{u} - \boldsymbol{\phi}(\mathbf{w}))$$

u

• By using the CEs, the latent solutions are approximated.

$$\mathbf{r} = \left(\sigma - \Omega_1 \sigma_0 - \Omega_2 \sigma_f - \Omega_3 \sigma'_0 - \Omega_4 \sigma'_f\right)^{\mathrm{T}} \beta_r + \Omega_1 \mathbf{r}_0 + \Omega_2 \mathbf{r}_f + \frac{\Omega_3 \mathbf{v}_0 + \Omega_4 \mathbf{v}_f}{b^2}$$
$$\lambda_r = \sigma^{\mathrm{T}} \beta_{\lambda_r}$$
$$\lambda_v = \sigma^{\mathrm{T}} \beta_{\lambda_v}$$
$$\mathbf{w} = \sigma^{\mathrm{T}} \beta_w$$
$$\nu = \sigma^{\mathrm{T}} \beta_{\nu}$$

• The first order necessary conditions and the corresponding losses are:

$$\dot{\mathbf{r}} = \frac{\partial H}{\partial \lambda_r} = \mathbf{v} \qquad \qquad \frac{\partial H}{\partial \mathbf{u}} = \mathbf{u} + \lambda_v + \mathbf{v} = \mathbf{0} \qquad \qquad \mathbf{\mathcal{L}}_a = \dot{\mathbf{v}} - \nabla U(\mathbf{r}) - \mathbf{M}\mathbf{v} - \phi(\mathbf{w}) \\ \mathbf{\mathcal{L}}_{\lambda_r} = \dot{\mathbf{\lambda}}_r + \nabla \nabla U^T \lambda_v \qquad \qquad \frac{\partial H}{\partial \mathbf{w}} = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \longrightarrow \mathbf{\mathcal{L}}_{\lambda_v} = \dot{\lambda}_v + \lambda_r + \mathbf{M}^T \lambda_v \\ \dot{\lambda}_r = -\frac{\partial H}{\partial \mathbf{r}} = -\nabla \nabla U^T \lambda_v \qquad \qquad \frac{\partial H}{\partial \mathbf{v}} = \mathbf{u} - \phi(\mathbf{w}) = \mathbf{0} \qquad \qquad \mathbf{\mathcal{L}}_u = \phi(\mathbf{w}) + \lambda_v + \mathbf{v} \\ \dot{\lambda}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \qquad \qquad \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \\ \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \qquad \qquad \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \qquad \qquad \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \\ \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \qquad \qquad \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \\ \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \qquad \qquad \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \\ \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \qquad \qquad \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0} \qquad \qquad \mathbf{\mathcal{L}}_w = 2\epsilon \mathbf{w} - \mathbf{v} \odot \phi'(\mathbf{w}) = \mathbf{0}$$

- The goal is to pass from a L1 Halo to a L2 Halo orbit.
- For PSO, the cost function associated to the particles is the sum of the time of flight and the norm of the loss vector.
- The other parameters used for this example are:

n	L	σ	$[u_{i,min}, u_{i,max}] \text{ km/s}^2$	с
100	35	Logistic	$[-5,5] \cdot 10^{-7}$	10

- A continuation procedure has been applied to decrease ٠ the value of ϵ from 1 to 10⁻⁶.
- The following boundary conditions are considered: •
 - $\mathbf{r}_0 = [0.823385182067467, 0, -0.022277556273235]$
 - $\mathbf{v}_0 = [0, 0.134184170262437, 0]$
 - $\mathbf{r}_f = [1.119601641146371, 0, 0.010405691913477]$
 - $\mathbf{v}_f = [0, 0.178317957880263, 0]$





1D fuel optimal lunar landing

- The 1D fuel optimal landing has been chosen to test the ability of the proposed framework to deal with possible discontinuities in the control.
- The fuel optimal problem is intrinsically constrained between a maximum and a minimum value. The use of a saturation function could result in a convenient choice.

 $\dot{h} = v$

$$\min \mathcal{J} = \int_{t_0}^{t_f} u(t) dt + \int_{t_0}^{t_f} \epsilon w^2(t) dt \quad \text{subject to:} \quad \begin{cases} \dot{v} = -g + u \\ t_0 \le t \le t_f \\ h(t_0) = h_0 \\ v(t_0) = v_0 \\ h(t_f) = h_f \\ v(t_f) = v_f \\ u_{\min} \le u \le u_{\max} \end{cases}$$

• The Hamiltonian of the problem is:

$$H = u + \lambda_h v + \lambda_v \left(-g + u\right) + \epsilon w^2 + \nu \left(u - \phi(w)\right)$$

1D fuel optimal lunar landing

• By using the CEs, the latent solutions are approximated.

$$h = \left(\boldsymbol{\sigma} - \Omega_{1}\boldsymbol{\sigma}_{0} - \Omega_{2}\boldsymbol{\sigma}_{f} - \Omega_{3}\boldsymbol{\sigma}_{0}' - \Omega_{4}\boldsymbol{\sigma}_{f}'\right)^{\mathrm{T}}\boldsymbol{\beta}_{h} + \Omega_{1}h_{0} + \Omega_{2}h_{f} + \frac{\Omega_{3}v_{0}}{b^{2}} + \frac{\Omega_{4}v_{f}}{b^{2}}$$
$$\lambda_{h} = \boldsymbol{\sigma}^{\mathrm{T}}\boldsymbol{\beta}_{\lambda_{h}}$$
$$\lambda_{v} = \boldsymbol{\sigma}^{\mathrm{T}}\boldsymbol{\beta}_{\lambda_{v}}$$
$$w = \boldsymbol{\sigma}^{\mathrm{T}}\boldsymbol{\beta}_{w}$$
$$\nu = \boldsymbol{\sigma}^{\mathrm{T}}\boldsymbol{\beta}_{\nu}$$

• The first order necessary conditions and the corresponding losses are:

$$\begin{split} \dot{h} &= \frac{\partial H}{\partial \lambda_h} = v & \frac{\partial H}{\partial u} = 1 + \lambda_v + \nu = 0 & \mathcal{L}_s &= \dot{v} + g - \phi(w) \\ \dot{v} &= \frac{\partial H}{\partial \lambda_v} = -g + u & \frac{\partial H}{\partial w} = 2\epsilon w - \nu \phi'(w) = 0 \longrightarrow \mathcal{L}_{\lambda_h} &= \dot{\lambda}_h \\ \dot{\lambda}_h &= -\frac{\partial H}{\partial h} = 0 & \frac{\partial H}{\partial w} = 2\epsilon w - \nu \phi'(w) = 0 \longrightarrow \mathcal{L}_{\lambda_v} &= \dot{\lambda}_v + \lambda_h \\ \dot{\lambda}_v &= -\frac{\partial H}{\partial v} = -\lambda_h & \frac{\partial H}{\partial \nu} &= u - \phi(w) = 0 & \mathcal{L}_w &= 2\epsilon w - \nu \phi'(w) \end{split}$$

1D fuel optimal lunar landing

- The following conditions are imposed: $r_0 = 10$ m, $v_0 = -2$ m/s, $t_f = 4.2$ s, g = 1.62 m/s².
- A continuation procedure has been applied to decrease both the value of ϵ from 100 to 10⁻¹⁵ and the coefficient c of the saturation function from c_0 to c_f .
- The parameters employed for the simulation are:



Conclusions

- A new PINN-based algorithm is proposed to solve optimal control problems with control constraints using saturation functions.
- X-TFC has been successfully employed to solve the TPBVP arising from the application of the PMP to the Hamiltonian of the problem.
- COCPs for space guidance have been solved
 - Minimum time-energy optimal Halo-Halo transfer.
 - 1D fuel optimal lunar landing. The bang-bang and bang-off-bang type of solution could be well approximated by means of saturation functions.
- The proposed algorithm can be potentially suitable for real time applications if a compromise between the optimality of the results and the precision of the dynamics is carried out.
- ➤ The possibility to solve more complex fuel optimal problems will be taken into account considering a combination of saturation functions.

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THANK YOU FOR YOUR ATTENTION!

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